



NASA's Physics of the Cosmos Program

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Interim PCOS Chief Scientist

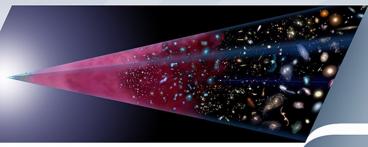
April APS 17 Apr 2021

Why Astrophysics?

Astrophysics is humankind's scientific endeavor to understand the universe and our place in it.



How did our universe begin and evolve?



Physics of the Cosmos (PCOS)



How did galaxies, stars, and planets come to be?



Cosmic Origins (COR)





Enduring National Strategic Drivers











Exoplanet
Exploration
(ExEP)

Program
Office
Themes



Physics of the Cosmos Program Office Purpose:

to explore some of the most fundamental questions regarding the physical forces and laws of the universe:

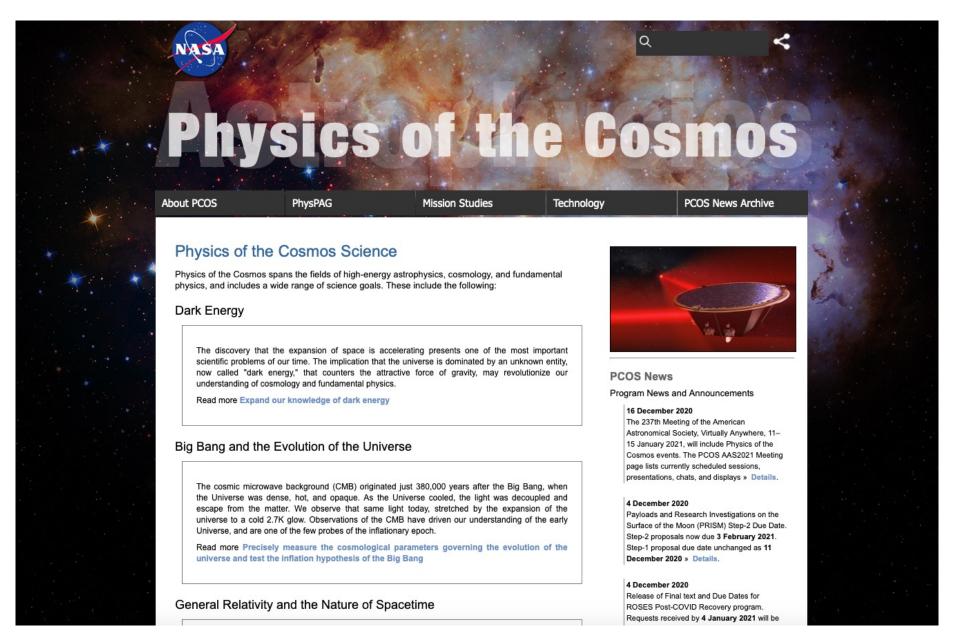
- the validity of Einstein's General Theory of Relativity and the nature of spacetime;
- the behavior of matter and energy in extreme environments;
- the cosmological parameters governing inflation and the evolution of the universe; and
- the nature of dark matter and dark energy.

Physics of the Cosmos spans the fields of high-energy astrophysics, cosmology, and fundamental physics, with a wide range of science goals. These include the following:

- General Relativity and the Nature of Spacetime
- Massive Black Holes and the Evolution of Galaxies
- Matter and Energy in the Most Extreme Environments
- Dark Energy
- Big Bang and the Evolution of the Universe

More resources: https://pcos.gsfc.nasa.gov

pcos.gsfc.nasa.gov





Activities supporting PCOS goals and priorities:

- Managed by the PCOS/COR Program Office at NASA's Goddard Space Flight Center and reported to NASA Headquarters.
- Include:
 - Mission studies and pre-project mission oversight, insight, and support
 - Strategic technology (SAT) maturation oversight, insight, and support
 - **Community engagement**, including via the Physics of the Cosmos Program Analysis Group (PhysPAG)
 - Maintaining **science cognizance** to enable more successful NASA strategic planning

The PCOS Program Office hosts

- ATHENA Study Office
- LISA Study Office

and oversees

- science and
- technology activities

for NASA's contribution to these ESA-led and other strategic missions.

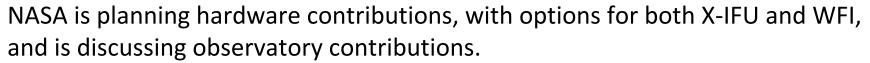


Athena

Athena is an ESA flagship X-ray mission slated for launch in the early 2030s

Two instruments provided by member states:

- calorimeter (X-IFU) and
- wide-field imager (WFI)



Current status: in (ESA) Phase B

- Nov 2019: Athena passed Mission Formulation Review (MFR)!
- Mission Adoption Review scheduled for June 2022
- US Athena Study will become a Project in May 2022 NASA Athena Science Team is co-chaired by Jon Miller (Michigan) and Laura Brenneman (SAO)

For more info:

http://www.the-athena-x-ray-observatory.eu/
https://asd.gsfc.nasa.gov/athena/





LISA



LISA is an ESA-led space gravitational wave observatory.

NASA is a junior partner with possible technology contributions, including:

- Laser (GSFC, JPL)
- Telescope (GSFC, U. Florida)
- Charge management system (U. Florida)

Additional NASA contributions will include

- Science Ground System Support
- Precursor Science programs
- Guest Investigator programs

Current status: in (ESA) Extended Phase A

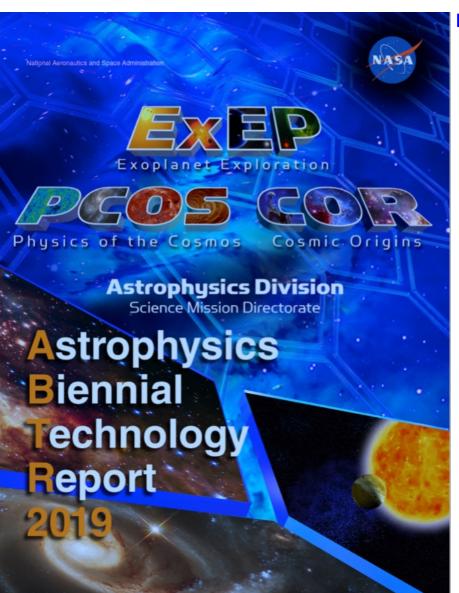
- Dec 2019: LISA passed ESA Mission Consolidation Review
- Mission Formulation Review in Fall 2021
- Mission Adoption currently anticipated in mid-2020s.

NASA LISA Study Team (Kelly Holley-Bockelmann, Chair) highlights:

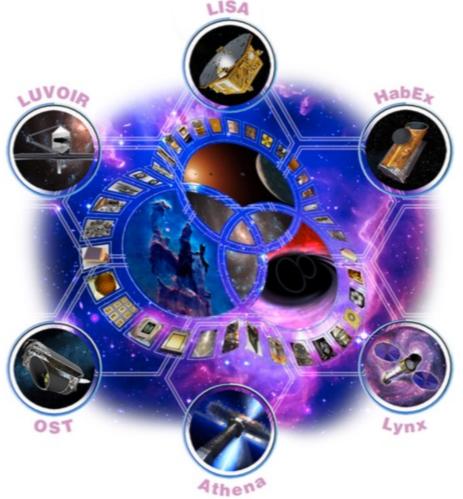
- Science Support Taskforce Report: Maximizing US Participation in LISA Science

For more info:

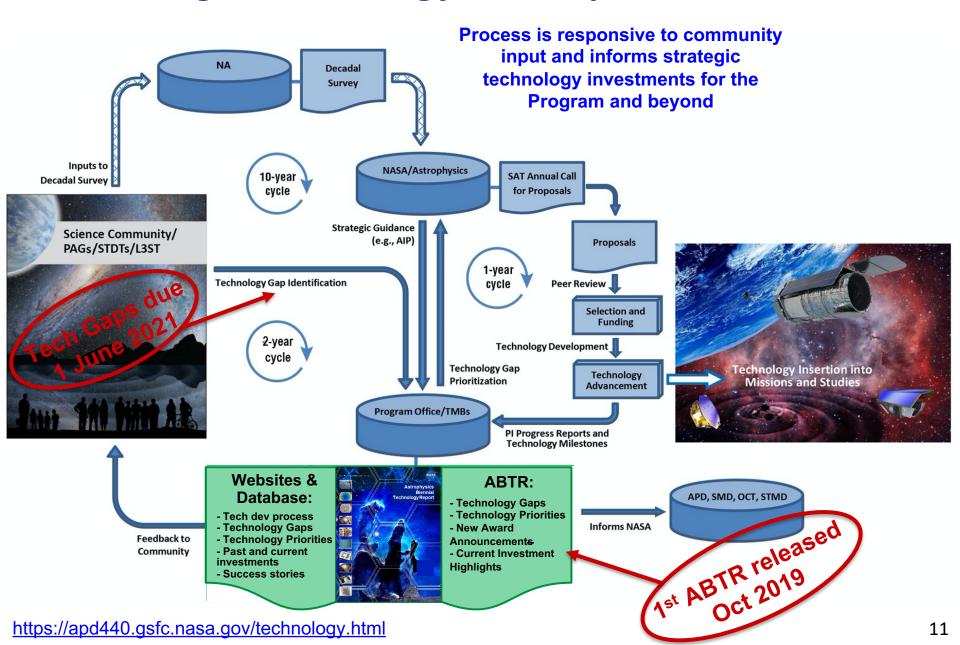
First Astrophysics Biennial Technology Report



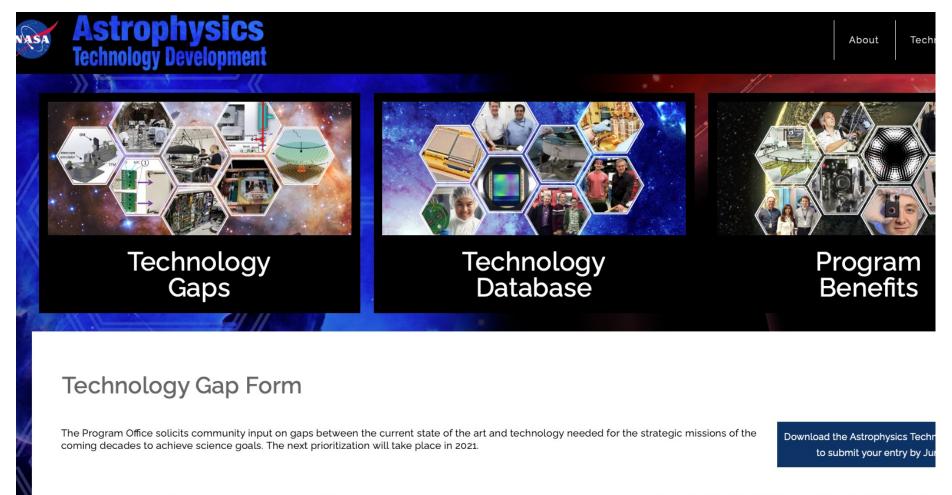
Process is responsive to community input and informs strategic technology investments for the Program and beyond



Strategic Technology Development Process



Technology Gap Solicitation



Due on June 1st OR 3 <u>weeks after the release of the</u>
<u>Astro2020 Decadal Survey</u>, whichever is *LATER*

Tier 1 Technology Gaps
Angular Resolution (UV/Vis/NIR)
Coronagraph Contrast
Coronagraph Contrast Stability
Cryogenic Readouts for Large-Format Far-IR Detectors
Fast, Low-Noise, Megapixel X-Ray Imaging Arrays with Moderate Spectral Resolution
High-Efficiency X-Ray Grating Arrays for High-Resolution Spectroscopy
High-Resolution, Large-Area, Lightweight X-Ray Optics
Large-Format, High-Resolution, UV/Vis Focal Plane Arrays
Large-Format, High-Spectral-Resolution, Small-Pixel X-Ray Focal-Plane Arrays
Large-Format, Low-Noise and Ultralow-Noise Far-IR Direct Detectors
Large-Format, Low-Noise, High-QE Far-UV Detectors
Next-Generation, Large-Format, Object Selection Technology for Multi-Object Spectrometers for LUVOIR

Vis/NIR Detection Sensitivity

Tier 1 Technology Gaps							
Angular Resolution (UV/Vis/NIR)							
Coronagraph	Tier 2 Technology Gaps						
Coronagraph	Advanced Millimeter-Wave Focal-Plane Arrays for CMB Polarimetry						
Cryogenic Rea	Detection Stability in Mid-IR						
	Heterodyne FIR Detector Arrays and Related Technologies						
Spectral Reso	High-Efficiency Object Selection Technology for UV Multi-Object Spectrometers						
High-Efficienc	High-Reflectivity Broadband FUV-to-NIR Mirror Coatings						
High-Resoluti							
Large-Format,							
Large-Format,							
Large-Format,	Spectrometers for HabEx						
Large-Format,	Starshade Deployment and Shape Stability						
Next-Generati	Starshade Starlight Suppression and Model Validation						
Spectrometers	Stellar Heriex Motion Sensitivity — Astrometry						
Vis/NIR Detec	Stellar Reflex Motion Sensitivity – Extreme Precision Radial Velocity						

Tier 1 Technology Gaps						
Angular Resolution (UV/Vis/NIR)						
Coronagraph	Tier 2 Technology Gaps					
Coronagraph	Advanced Millimeter-Wave Focal-Plane Arrays for CMB Polarimetry					
Cryogenic Rea	Detection Stabil	ection Stabil Tier 3 Technology Gaps				
Fast, Low-Noi		Advanced Cryocoolers				
Spectral Reso	High-Efficiency	5.5				
High-Efficienc	mign-renonnan	Large Cryogenic Optics for the Mid-IR to Far-IR				
High-Resoluti	High-Reflectivity	g-Wavelength-Blocking Filters for X-Ray Micro-Calorimeters r-Noise, High-QE UV Detectors				
Large-Format, Large-Format,	High-Inroughpi					
	Large-Format O	Low-Stress, Highly Stable X-Ray Reflective Coatings				
	Spectrometers for Starshade Deplo	Photon-Counting, Large-Format UV Detectors				
	Starshade Starli	Polarization-Preserving Millimeter-Wave Optical Elements				
Spectrometers		UV Coatings				
Vis/NIR Detec		UV Detection Sensitivity				
	Sicilal Helick IVI	UV/Vis/NIR Tunable Narrow-Band Imaging Capability				
Warm Readout Electronics for Large-Format Far-IR Detectors						

	Tie	r 1 Technology Gaps	•				
Angular Resol	ution (UV/Vis/NI						
Coronagraph	Tier 2 Technology Gaps						
Coronagraph	Advanced Millimeter-Wave Focal-Plane Arrays for CMB Polarimetry						
Cryogenic Rea	Detection Stabil	Stabil Tier 3 Technology Gaps					
Fast, Low-Noi	Heterodyne FIR	rodyne FIR Advanced Cryocoolers					
Spectral Reso	High-Efficiency	High-Perform	Tier 4 Technology Ga	ps			
High-Efficienc	riigir-i ciloimaii	Large Cryoge Compact, Integ	grated Spectrometers for 100 to 1000	μm			
mgii moodian	High-Reflectivity	Low-Noise, F Rapid Readout Electronics for X-Ray Detectors O Low-Stress Short Ways Ltd Contings					
Larye-i ormai,	High-Throughpo						
Large-Format,	Large-Format O						
	Spectrometers for Starshade Deplo	I Hotoli Couliting, Large I of	Tier 5 T	echnology Gaps			
	Starshade Starli	Folalization-Fleserving willi	Advancement of X-Ray Polarimeter S	Sensitivity			
Spectrometers	Stellar Reflex M	UV Coalings	Far-IR Spatio-Spectral Interferometry				
Vis/NIR Detec		UV Detection Sensitivity	High-Precision Low-Frequency Radi	io Spectrometers and Interferome	ters		
	Stellar Helick IVI	UV/Vis/NIR Tunable Narrow-	Mid-IR Coronagraph Contrast				
Warm Readout Electronics fo							
			Very-Wide-Field Focusing Instrument for Time-Domain X-Ray Astronomy				
			Wide-Bandwidth, High-Spectral-Dyr Radio-Frequency Observations on the		or Low-		



PCOS Chief Scientist enables ground-breaking science from space by working at the interfaces between missions and studies, technology, the community, and NASA HQ.

Current PCOS Science Goals and Priorities:

- Ensure a more successful **Decadal survey** by supporting community
 preparations and HQ activities, spanning the range of inputs: from science to
 missions, technology, and state of the profession, which all impact our ability
 to do ground-breaking science
- Ensure more **successful missions** by
 - supporting ongoing mission studies and pre-projects, e.g., LISA, Lynx, Athena
 - through technology efforts, e.g., SAT;
 - by coordinating with current missions; and
 - by preparing for studies for mission recommended by the Astro2020 Decadal
- **Engage the community** to support a successful APD portfolio.

PhysPAG

Physics of the Cosmos Program Analysis Group

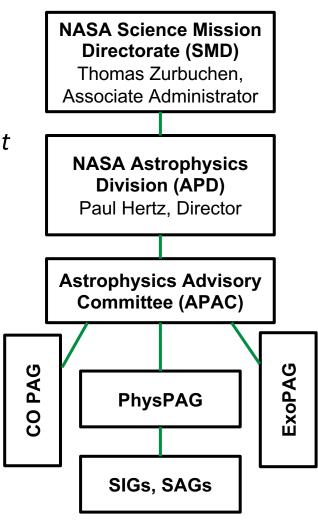
- Purpose:

- provide input to NASA relevant to PCOS
- help NASA inform interested parties about PCOS doings
- Membership: You!

Anyone interested in providing input to NASA relevant to its Physics of the Cosmos Program

- Leadership:
 - **Executive Committee** (EC):
 - Chair Emeritus: Graça Rocha
 - Chair: Ryan Hickox
 - Vice Chair: Grant Tremblay
 - 11 EC members chair 6 Science Interest Groups (SIGs): longer-standing discipline-specific
 - support formation of Science Analysis Groups
 (SAGs): group created to analyze a specific science question
 - facilitate info flow between NASA and community

Communication Network:







Let's Chat!

Sparks

- What have you found useful?
 - SIGs? SAGs?
 - professional exchange of ideas?
 - white paper preparations?
 - Strategic Astrophysics Technology (SAT) program?
 - Î
- What would you like to see more of? or less of?
 - more community leadership?
 - [
- What do you need from NASA?
- What are you concerned about?

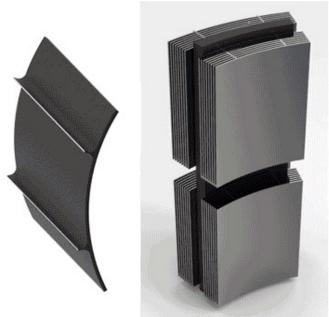




Conclusions



- PCOS will continue to work with the other program offices (COR, ExEP) to ensure that NASA is fully prepared for any recommendations made by the Astro2020 Decadal Survey
- PCOS will continue to provide insight and support for mission studies in development
- PCOS will continue to oversee the development of strategic technologies necessary for astrophysics in the coming years and decades
- PCOS, via the PhysPAG, will continue to offer scientific insight to NASA on various topics



Single mirror segment plus stacked segments with support panel PI: Zhang, William (GSFC)

Sign up for our mailing list!

https://pcos.gsfc.nasa.gov/pcosnews-mailing-list.php